SuperB/ http://www.pi.infn.it/SuperB/

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CIPANP '09, San Diego, May 2009.





Overview

- SuperB in a nutshell
- Physics potential of SuperB
 - New Physics Search Capabilities
 - Lepton Flavour & CP Violation in τ decay
 - Rare Decays
 - Fundamental Symmetries & Higgs/dark matter
 - Standard Model measurements.
- Current Status
- Conclusion



SuperB (In a Nutshell)

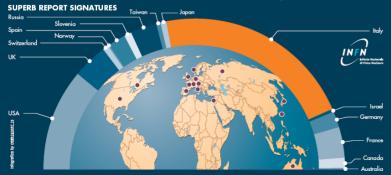
Site: Tor Vergata Campus (Rome II)

- Asymmetric energy e⁺e⁻ collider
- Low emittance operation (like LC)
- Polarised beams
- Luminosity 10³⁶ cm⁻²s⁻¹
 - 75ab⁻¹ data at the Y(4S)
 - Collect data at other √s
 - Start data taking as early as 2015





International Community



Geographical distribution of CDR signatories.

Precision B, D and τ decay studies and spectroscopy

- New Physics in loops
 - 10 TeV reach at 75ab⁻¹
 - Rare decays
 - △S CP violation measurements
- Lepton Flavour & CP Violation in τ decay
- Light Higgs searches
- Dark Matter searches

http://www.pi.infn.it/SuperB/



Physics Potential of SuperB

- SuperB is a super flavour factory!
 - Hundreds of billions of B_u, B_d, D, τ decays with 5 years of data taking:
 - 75ab⁻¹ at the Υ(4S).
 - Running at charm threshold: ψ(3770).
 - And at other Y resonances.
- Lepton Flavour & CP Violation in τ decay.
- Rare Decay constraints on new physics.
- Fundamental symmetries & Higgs/dark matter constraints:
 - Lepton Universality, CP, T, CPT.
- + Standard Model B, D and τ physics.



Lepton Flavor Violation in τ decay

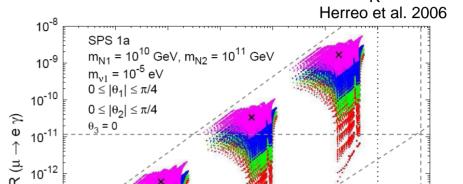
10⁻¹³

10⁻¹⁴

■ $\tau \rightarrow \mu \gamma$ upper limit can be correlated to θ_{13} (neutrino mixing/CPV, T2K etc.) and also to $\mu \rightarrow e \gamma$.

SUSY seasaw = CMSSM + $3\nu_R$ + ∇

- Complementary to flavour mixing in quarks.
- Golden modes:
 - $\tau \rightarrow \mu \gamma$ and 3μ .
- e- beam polarization:
 - Lower background
 - Better sensitivity than competition!
- e+ polarization used later in programme.
- CPV in $\tau \rightarrow K_S \pi \nu$ at the level of ~10⁻⁵.
- Bonus:
 - Can also measure τ g-2 (polarization is crucial).
 - $\sigma(g-2) \sim 2.4 \times 10^{-6}$ (statistically dominated error).



Process	Expected 90%CL	4σ Discovery
	upper limited	Reach
$\mathcal{B}(au o \mu \gamma)$	2×10^{-9}	5×10^{-9}
$\mathcal{B}(au o \mu \mu \mu)$	2×10^{-10}	8.8×10^{-10}

BR $(\tau \rightarrow \mu \gamma)$

 $m_{N3} = 10^{13} \text{ GeV}$

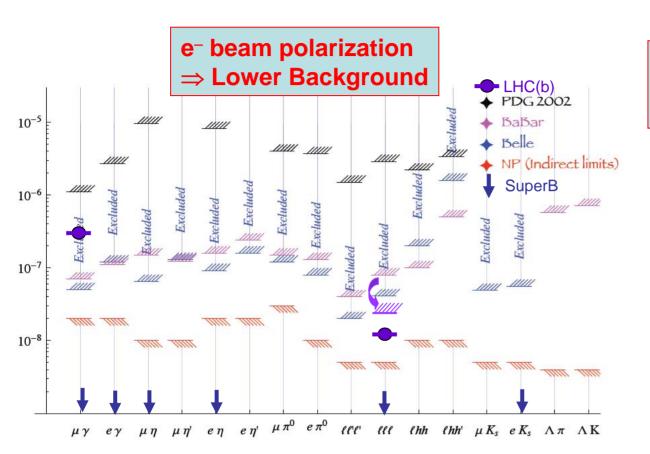
10⁻¹⁰

Use $\mu \gamma/3I$ to distinguish SUSY vs. LHT.

5



Lepton Flavor Violation in τ decay



SuperB Sensitivity (75ab⁻¹)

Process	Sensitivity
$\mathcal{B}(\tau \to \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \to e \gamma)$	2×10^{-9}
$\mathcal{B}(au o \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \to eee)$	2×10^{-10}
$\mathcal{B}(\tau \to \mu \eta)$	4×10^{-10}
$\mathcal{B}(\tau \to e \eta)$	6×10^{-10}
$\mathcal{B}(au o \ell K_s^0)$	2×10^{-10}

- LHC is not competitive (Re: both GPDs and LHCb).
- SuperB sensitivity ~10 50× better than NP allowed branching fractions.



Rare Decays

No one smoking gun... rather a 'golden matrix'.

X = Golden Channel	H^+	Minimal	Non-Minimal	Non-Minimal	NP	Right-Handed
o = Observable effect	high $tan\beta$	FV	FV (1-3)	FV (2-3)	Z-penguins	currents
$BR(B \to X_s \gamma)$		X		О		О
$A_{CP}(B \to X_s \gamma)$				X		О
${ m BR}(B o au u)$	X-CKM					
$BR(B \to X_s l^+ l^-$)			O	О	О
$BR(B \to K \nu \overline{\nu})$				О	X	
$S(K_S\pi^0\gamma)$						X
β (Δ S)			X- CKM			X

 Need to measure all observables in order to select/eliminate new physics scenarios!

Mode	Sensitivity		
	Current	$10 {\rm ~ab^{-1}}$	75 ab^{-1}
$\mathcal{B}(B \to X_s \gamma)$	7%	5%	3%
$A_{CP}(B o X_s \gamma)$	0.037	0.01	0.004 – 0.005
$\mathcal{B}(B^+ \to au^+ u)$	30%	10%	3-4%
$\mathcal{B}(B^+ \to \mu^+ \nu)$	X	20%	56%
$\mathcal{B}(B \to X_s l^+ l^-)$	23%	15%	46%
$A_{\rm FB}(B \to X_s l^+ l^-)_{s_0}$	X	30%	4-6%
$\mathcal{B}(B \to K \nu \overline{\nu})$	X	X	1620%
$S(K_S^0\pi^0\gamma)$	0.24	0.08	0.02 – 0.03

- •The golden modes
 - will be measured by SuperB.
 - `smoking guns' for their models.
- •Measurements not yet made are denoted by X.
- •With 75ab⁻¹ we can
 - Reach above a TeV with B→ τν
 - See B→Kνν



Rare Decays: SUSY CKM

Flavour couplings in squark sector (like CKM)

$$\text{M}^{2}\text{a} \approx \begin{pmatrix} m_{\tilde{d}_{L}}^{2} & m_{d}(A_{d} - \mu \tan \beta) & (\Delta_{12}^{d})_{LL} & (\Delta_{12}^{d})_{LR} & (\Delta_{13}^{d})_{LL} & (\Delta_{13}^{d})_{LL} & (\Delta_{13}^{d})_{LR} \\ L_{HC} & m_{\tilde{d}_{R}}^{2} & (\Delta_{12}^{d})_{RL} & (\Delta_{12}^{d})_{RR} & (\Delta_{13}^{d})_{RL} & (\Delta_{13}^{d})_{RL} & (\Delta_{13}^{d})_{RR} \\ & & m_{\tilde{s}_{L}}^{2} & m_{\tilde{s}}(A_{s} - \mu \tan \beta) & (\Delta_{23}^{d})_{LL} & (\Delta_{23}^{d})_{LR} \\ & & & m_{\tilde{b}_{L}}^{2} & m_{b}(A_{b} - \mu \tan \beta) \\ & & & m_{\tilde{b}_{L}}^{2} & m_{b}(A_{b} - \mu \tan \beta) \end{pmatrix}$$

and similarly for M^2

- NP scale: m_q
- Flavour and CP violation coupling: $(\Delta_{ij}^d)_{AB}/m_{\tilde{q}}^2$
- Why?
 - Non trivial CKM & MSW, so it is natural for squarks to mix!
 - Unnatural to have couplings ~0 and a low mass scale.
- e.g. MSSM: 124 parameters (160 with v_R).
 - Most are flavour couplings!



Rare Decays: SUSY CKM

- Couplings are $\left(\delta_{ij}^{q}\right)_{AB} = \left(\Delta_{ij}^{q}\right)_{AB} / m_{\tilde{q}}^{2}$ where A, B=L, R, and i, j are squark generations.
- e.g. Constraint on $(\delta_{23}^{d})_{LR}$ using:

•
$$\mathcal{B}(B \rightarrow X_s \gamma)$$
 [green]

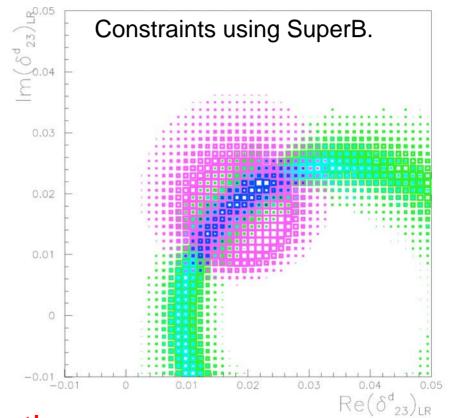
- $\mathcal{B}(B \rightarrow X_s l^+ l^-)$ [cyan]
- $A_{CP}(B \rightarrow X_s \gamma)$ [magenta]
- Combined [blue]

SuperB probes new physics in SUSY larger than 20TeV (and up to 300TeV in some scenarios)

Now: Whole Range allowed.

SuperB: % level measurement!

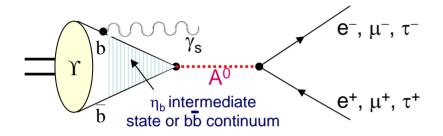
See (or rule out) a sparticle signal at TeV scale with 75ab⁻¹.





Searching for a Light Higgs & Dark Matter

- Many NP scenarios have a possible light Higgs Boson (e.g. 2HDM).
- Can use $Y(nS) \rightarrow l^+l^-$ to search for this.
 - Contribution from A⁰ would break lepton universality



M. A. Sanchis-Lozano, hep-ph/0510374, Int. J. Mod. Phys. A19 (2004) 2183

- Can expect to record at least 300fb⁻¹ recorded at the Υ(3S) in SuperB.
- This is $10 \times$ the BaBar data sample at the $\Upsilon(3S)$.

NMSSM Model with 7 Higgs Bosons

Physical Higgs bosons: (seven)

2 neutral CP-odd Higgs bosons (A_{1,2}) 3 neutral CP-even Higgs bosons (H_{1,2,3}) 2 charged Higgs bosons (H±)

A₁ could be a light DM candidate.

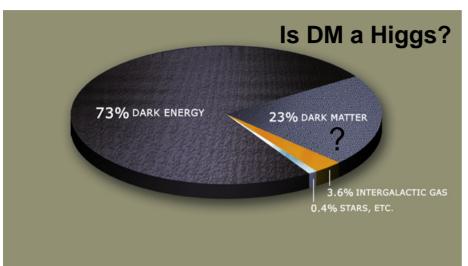
Possible NMSSM Scenario

 $A_1 \sim 10 \text{ GeV}$ $H_1 \sim 100 \text{ GeV (SM-like)}$ Others ~300 GeV (almost degenerate)

Gunion, Hooper, McElrath [hep-ph:0509024] McElrath [hep-ph/0506151], [arXiv:0712.0016]



Searching for Dark Matter

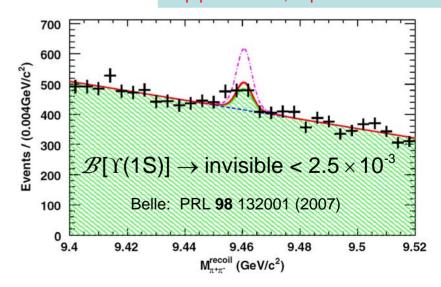


 Possible to search for the effect of DM at the B-factories for most modes:

$$\begin{array}{ccc} \Upsilon \rightarrow invisible & J/\Psi \rightarrow invisible \\ \eta \rightarrow invisible & \Upsilon \rightarrow \gamma + invisible \\ B^+ \rightarrow K^+ + invisible & \Upsilon \rightarrow \gamma A_1, A_1 \rightarrow \tau^+ \tau^- \\ K^+ \rightarrow \pi^+ + invisible & J/\Psi \rightarrow \gamma A_1 \end{array}$$

hep-ph/0506151, hep-ph/0509024, hep-ph/0401195, hep-ph/0601090, hep-ph/0509024, hep-ex/0403036 ...

- SM Expectation: $\mathcal{B}(\Upsilon(1S) \rightarrow \nu \overline{\nu}) = (9.9 \pm 0.5) \times 10^{-6}$
- NP extension: $\mathcal{B}(\Upsilon(1S) \to \chi \chi)$ up to 6×10^{-3}
- SuperB should be able to provide a precision constraint on this channel.
- Belle has 7fb-1 at the Υ(1S), SuperB will have hundreds of fb⁻¹.

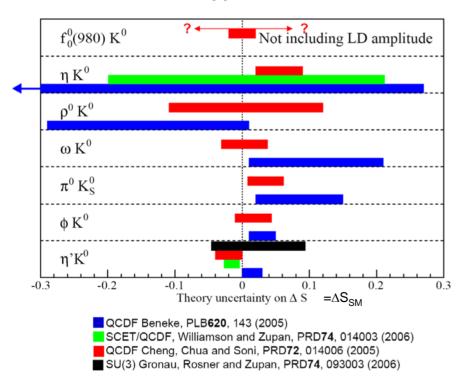


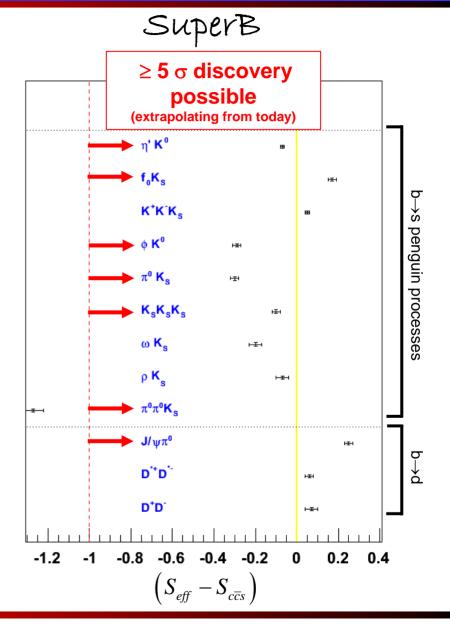


ΔS measurements (CPV)

- β=(21.1±0.9)° from Charmonium decays.
- Look in many different b→s and b→d decays for sin2β deviations from the SM:

$$\Delta S_{\rm NP} = S_{eff} - S_{c\overline{c}s} - \Delta S_{\rm SM}$$







Standard Model measurements.

B Physics at Y(4S)

	Observable	B Factories (2 ab ⁻¹)	SuperB (75 ab
	$\sin(2\beta) (J/\psi K^0)$	0.018	0.005 (†)
	$cos(2\beta) (J/\psi K^{*0})$	0.30	0.05
	$sin(2\beta)$ (Dh^0)	0.10	0.02
	$cos(2\beta)$ (Dh ⁰)	0.20	0.04
	$S(J/\psi \pi^0)$	0.10	0.02
	$S(D^{+}D^{-})$	0.20	0.03
	$S(\phi K^0)$	0.13	0.02(*)
	$S(\eta'K^0)$	0.05	0.01(*)
	$S(K_s^0K_s^0K_s^0)$	0.15	0.02 (*)
	$S(K_s^0\pi^0)$	0.15	0.02(*)
	$S(\omega K_s^0)$	0.17	0.03(*)
	$S(f_0K_s^0)$	0.12	0.02 (*)
	γ (B \rightarrow DK, D \rightarrow CP eigenstate	s) $\sim 15^{\circ}$	2.5°
	γ (B \rightarrow DK, D \rightarrow suppressed sta	ates) $\sim 12^{\circ}$	2.0°
	γ ($B \rightarrow DK$, $D \rightarrow$ multibody sta	ites) $\sim 9^{\circ}$	1.5°
	γ (B \rightarrow DK, combined)	$\sim 6^{\circ}$	$1-2^{\circ}$
	$\alpha \ (B \to \pi \pi)$	$\sim 16^{\circ}$	3°
	$\alpha (B \rightarrow \rho \rho)$	$\sim 7^{\circ}$	1−2° (∗)
	$\alpha \ (B \rightarrow \rho \pi)$	$\sim 12^{\circ}$	2°
	α (combined)	$\sim 6^{\circ}$	$1-2^{\circ}$ (*)
	$2\beta + \gamma \left(D^{(\star)\pm}\pi^{\mp}, D^{\pm}K_{s}^{0}\pi^{\mp}\right)$	20°	5°
	$ V_{cb} $ (exclusive)	4% (*)	1.0% (*)
	$ V_{cb} $ (inclusive)	1% (*)	0.5% (*)
	$ V_{ub} $ (exclusive)	8% (*)	3.0% (*)
	$ V_{ub} $ (inclusive)	8% (*)	2.0% (*)
		576 (1)	21070 (1)
	$B(B \rightarrow \tau \nu)$	20%	4% (†)
	$B(B \rightarrow \mu\nu)$	visible	5%
	$\mathcal{B}(B \to D\tau\nu)$	10%	2%
	$\mathcal{D}(D \to D T V)$	10/0	270
	$B(B \rightarrow \rho \gamma)$	15%	3% (†)
	$\mathcal{B}(B \to \omega \gamma)$	30%	5%
	$A_{CP}(B \rightarrow K^*\gamma)$		
ĺ	- (0.007 (†)	0.004 († *)
	$A_{CP}(B \to \rho \gamma)$	~ 0.20	0.05
	$A_{CP}(b \rightarrow s\gamma)$	0.012 (†)	$0.004 (\dagger)$
ĺ	$A_{CP}(b \rightarrow (s + d)\gamma)$	0.03	0.006 (†)
	$S(K_S^0\pi^0\gamma)$	0.15	0.02 (*)
	$S(\rho^0\gamma)$	possible	0.10
	. ,		
	$A_{CP}(B \rightarrow K^*\ell\ell)$	7%	1%
	$A^{FB}(B \rightarrow K^*\ell\ell)s_0$	25%	9%
ı	$A^{FB}(B \rightarrow X_s \ell \ell) s_0$	35%	5%
IJ	$\mathcal{B}(B \to K \nu \overline{\nu})$	visible	20%
1	$B(B \rightarrow K \nu \bar{\nu})$ $B(B \rightarrow \pi \nu \bar{\nu})$	VIDIDIC	possible
	$D(D \rightarrow \pi \nu \nu)$	_	possible

Rare Charm Decays: 1 month at ψ(3770)

Channel	Sensitivity
$D^0 \to e^+ e^-, D^0 \to \mu^+ \mu^-$	1×10^{-8}
$D^0 \to \pi^0 e^+ e^-, \ D^0 \to \pi^0 \mu^+ \mu^-$	2×10^{-8}
$D^0 \rightarrow \eta e^+ e^-, D^0 \rightarrow \eta \mu^+ \mu^-$	3×10^{-8}
$D^0 \to K_s^0 e^+ e^-, D^0 \to K_s^0 \mu^+ \mu^-$	3×10^{-8}
$D^+ \to \pi^+ e^+ e^-, \ D^+ \to \pi^+ \mu^+ \mu^-$	1×10^{-8}
$D^0 \to e^{\pm} \mu^{\mp}$	1×10^{-8}
$D^+ \to \pi^+ e^{\pm} \mu^{\mp}$	1×10^{-8}
$D^0 \to \pi^0 e^\pm \mu^\mp$	2×10^{-8}
$D^0 \to \eta e^{\pm} \mu^{\mp}$	3×10^{-8}
$D^0 \rightarrow K_s^0 e^{\pm} \mu^{\mp}$	3×10^{-8}
-	
$D^+ \to \pi^- e^+ e^+, D^+ \to K^- e^+ e^+$	1×10^{-8}
$D^+ \to \pi^- \mu^+ \mu^+, \ D^+ \to K^- \mu^+ \mu^+$	1×10^{-8}
$D^+ \to \pi^- e^{\pm} \mu^{\mp}, D^+ \to K^- e^{\pm} \mu^{\mp}$	1×10^{-8}

τ: LFV / CPV / ...

Process	Sensitivity
$B(\tau \rightarrow \mu \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \to e \gamma)$	2×10^{-9}
$\mathcal{B}(\tau \to \mu \mu \mu)$	2×10^{-10}
$\mathcal{B}(\tau \to eee)$	2×10^{-10}
$B(\tau \rightarrow \mu \eta)$	4×10^{-10}
$B(\tau \rightarrow e \eta)$	6×10^{-10}
$\mathcal{B}(au o \ell K_s^0)$	2×10^{-10}

Mode	Observable	$\Upsilon(4S)$	$\psi(3770)$	\circ
<u></u>		(75 ab^{-1})	(300 fb^{-1})	D
$D^0 \rightarrow K^+\pi^-$	x'^2	3×10^{-5}		arm
	y'	7×10^{-4}		3
$D^0 \rightarrow K^+K^-$	y_{CP}	5×10^{-4}		
$D^0 \rightarrow K_S^0 \pi^+ \pi^-$	x	4.9×10^{-4}		\leq
	y	3.5×10^{-4}		×.
	q/p	3×10^{-2}		<u> </u>
	ϕ	2°		<u></u>
$\psi(3770) \rightarrow D^0 \overline{D}^0$	x^2		$(1-2) \times 10^{-5}$	<u>u</u>
	y		$(1-2) \times 10^{-3}$	
	$\cos \delta$		(0.01-0.02)	

See CDR and Valencia report for details of the SM measurements and other possible NP searches.

B Physics at Y(5S)

Observable	Error with 1 ab^{-1}	Error with 30 ab^{-1}
ΔΓ	$0.16 \ \mathrm{ps^{-1}}$	$0.03~{\rm ps}^{-1}$
Γ	$0.07~{\rm ps}^{-1}$	$0.01~{\rm ps^{-1}}$
β_s from angular analysis	20°	8°
$A^s_{ m SL}$	0.006	0.004
A_{CH}	0.004	0.004
$\mathcal{B}(B_s \to \mu^+ \mu^-)$	-	$<8\times10^{-9}$
$ V_{td}/V_{ts} $	0.08	0.017
$\mathcal{B}(B_s \to \gamma \gamma)$	38%	7%
β_s from $J/\psi\phi$	10°	3°
$\beta_s \text{ from } B_s \to K^0 \bar{K}^0$	24°	11°

Adrian Bevan 13



Current Status

- Work toward realizing the experiment is starting!
 - Prepare Technical Design Reports now end 2010.
 - Construct experiment 2011-2015.

... then start to take data!

- Just started work toward Technical Design Reports.
 - Fast and GEANT simulations available to optimize detector studies for physics goals!
 - Detector and accelerator R&D ongoing.
 - Preparing
 - A Technical Proposal for end of 2009.
 - A Technical Design Report for end of 2010/start 2011.
 - Physics workshop at Warwick April 2009.
 - Plenty of work to do…
- You're welcome to join the effort!
 - contact giorgi.superb@pi.infn.it for more information.



Conclusion

- SuperB measures a unique set of new physics observables to high precision.
 - Measure a golden matrix to test new physics signatures.

Scenario 1

- LHC finds new physics after 2010...
 - But what exactly has it found...? ⇒ <u>SuperB pins it down!</u>

May 2009 Adrian Bevan 15



Conclusion

- SuperB measures a unique set of new physics observables to high precision.
 - Measure a golden matrix to test new physics signatures.

Scenario 1

- LHC finds new physics after 2010...
 - But what exactly has it found...? ⇒ <u>SuperB pins it down!</u>

Scenario II

- LHC doesn't find new physics!
 - SuperB indirectly searches far beyond the reach of LHC: ~ 100TeV.
 - c.f. 1970: GIM mechanism and K₁→μ⁺μ⁻.
 - c.f. 1973: 3rd generation from CKM matrix.
 - c.f. 90s \Rightarrow heavy top quark mass from Δm_d .
 - Z⁰ width constraints on the SM Higgs ...?



Conclusion

- SuperB measures a unique set of new physics observables to high precision.
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Scenario 1

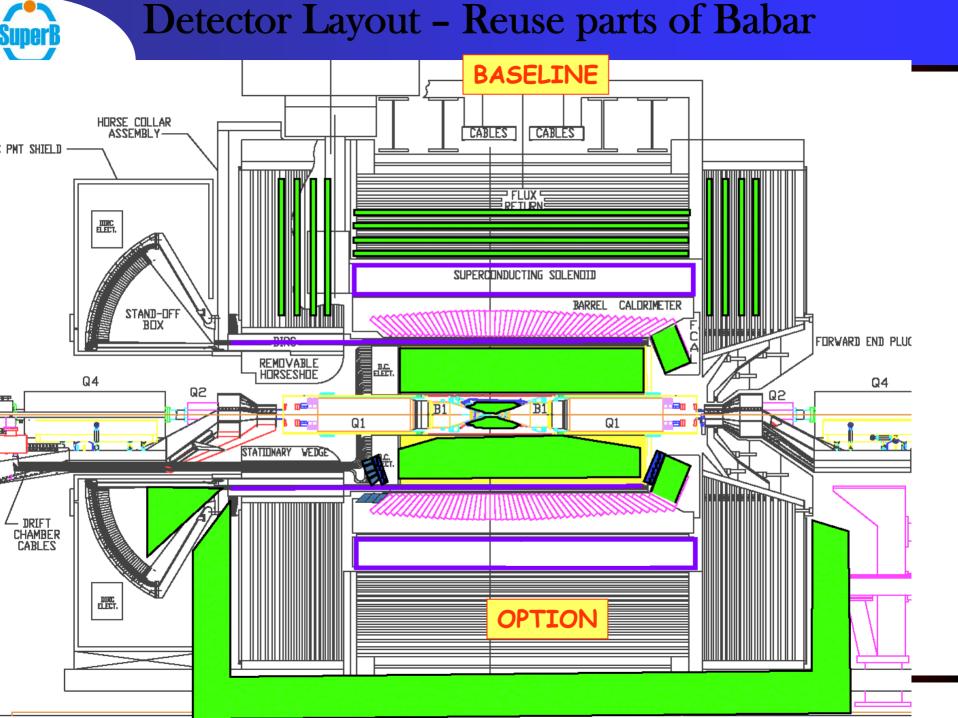
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Scenario II

- LHC doesn't find new physics!
 - ⇒SuperB indirectly searches far beyond the reach of LHC: ~ 100TeV.
 - c.f. 1970: GIM mechanism and $K_1 \rightarrow \mu^+\mu^-$.
 - c.f. 1973: 3rd generation from CKM matrix.
 - c.f. 90s ⇒ heavy top quark mass from ∆m_d.
 - Z⁰ width constraints on the SM Higgs ...?
- Either way round we need SuperB to unravel the mysteries of nature.
- + SuperB does all of the Standard Model measurements you'd expect.
 - and ... constrains SM errors for other NP searches; e.g. $K^0_L \rightarrow \pi^0 \nu \bar{\nu}$.
- See http://www.pi.infn.it/SuperB/, CDR, and physics workshop report for more details: arXiv:0810.1312 and 0709.0451.



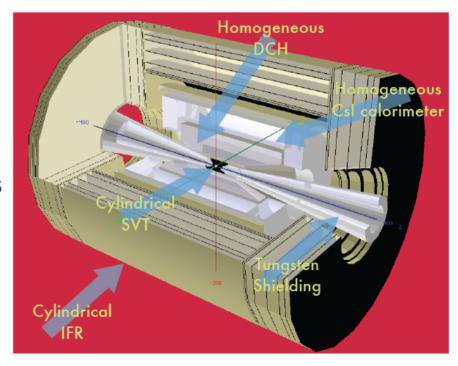
Additional Material



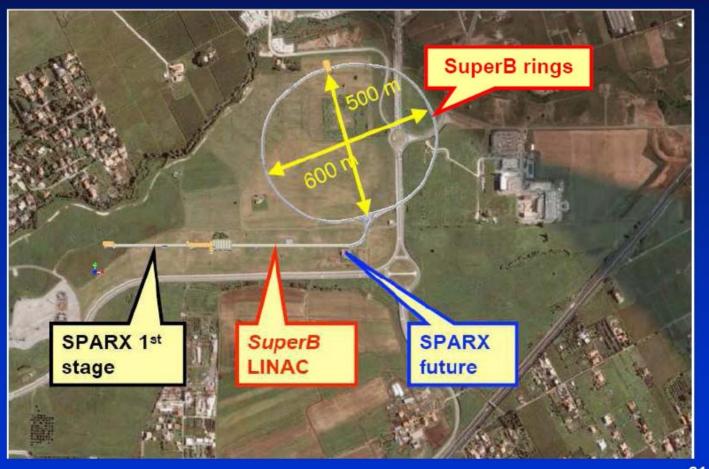


Detector

- Simulation:
 - FastSim (validated on using geometry for BaBar)
 - Reproduces BaBar resolutions etc.
 - Change to SuperB geometry and boost for development of benchmark studies.
 - Then move to GEANT 4 for more detailed work.
 - GEANT 4 model of SuperB shown.
 - Using BaBar framework.
 - Draw on a decade of analysis experience from BaBar and Belle to optimize an already good design.



SuperB footprint on Tor Vergata site

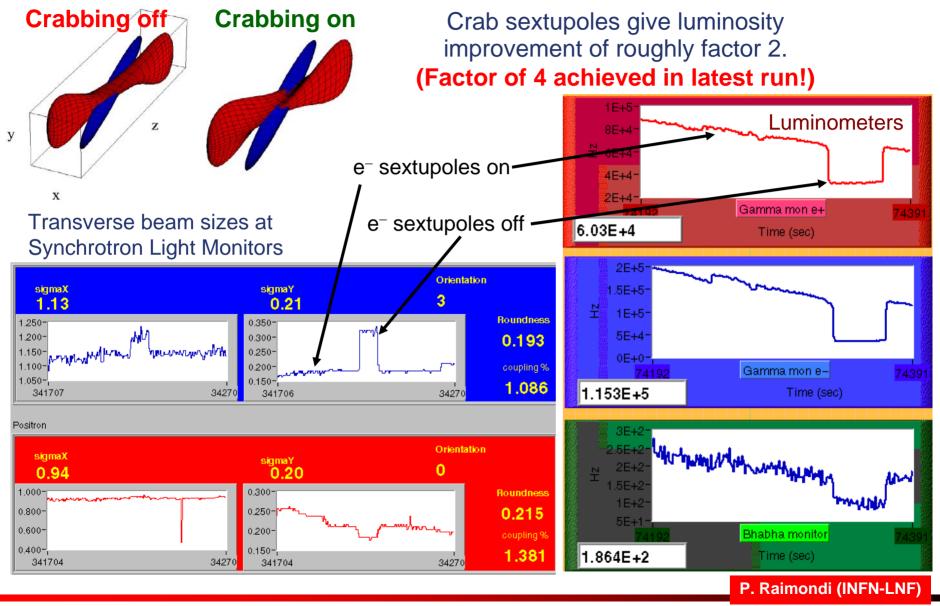




21



Crab waist tests at DAФNE



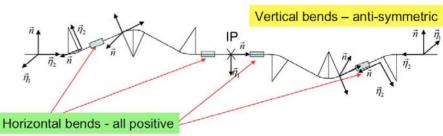


Polarisation

- In a storage ring, particle spins naturally precess around the vertical fields of the arc dipoles, at a rate determined by the particle energy.
 - This means that vertical polarisation is naturally preserved, but longitudinal polarisation can be lost without preventive measures.

For SuperB, there are two options to maintain longitudinal polarisation in the beam at the IP:

- 1. Use solenoids opposite the IP, to rotate the spin by π around the longitudinal axis.
- 2. Use solenoids or vertical bends to rotate between vertical and longitudinal spin before and after the IP.
- Option 2 will probably work best for the multi-GeV SuperB rings, but more studies are needed.
- With a source providing ~ 90%
 polarisation, it is expected that an average polarisation of 80% can be achieved for the e⁻ beam. e⁺ polarisation would be an upgrade...



Snake

IP